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in a platinum crucible, the oxid of mercury was decomposed, and its elements expelled, and a small portion of a green oxid remained in the crucible. In several repetitions of the process this, invariably, occurred. I had been led to suspect that this was the oxid of nickel, because the alkaline solution, from which it had been obtained, gave a black precipitate, with the hidro-sulphuret of ammonia; accordingly, on fusing a portion of this oxid, with borax, under the blow-pipe, it produced a glass of a hyacinth red; the same fact took place with a portion of a substance, known to be the oxid of nickel, which was fused with borax, for the sake of comparison. On fusing a portion of the chromat of lead, or Siberian red lead ore, with borax, and afterwards with vitreous phosphoric acid, glasses of an emerald green colour were produced.

Hence it was concluded, that the meteoric stones of Weston do not contain *chrome*, but that the green oxid obtained, was the oxid of nickel.

#### No. LII.

Observations of the Comet which appeared in September 1807, in the Island of Cuba, by J. J. de Ferrer.

Read August 19th, 1808.

		Mean time at the City of Havanna.	The observed long. of the Comet.	The observed lat. of the Comet.
1807. Octr.	1	6 54 50	220 21 12	18 46 03 N.
	18		234 36 58	37 41 11
Novr.	3	6 56 05	251 41 25	51 13 00
	4	6 49 30	252 57 08	51 54 42
	7	6 44 20	257 02 22	53 54 1 <b>8</b>
	17	7 04 26	272 54 40	<b>59 17 31</b>
	18	6 27 36	274 37 42	59 42 37
	19	6 44 10	276 27 40	60 06 13
	25	6 59 07	287 53 57	61 56 32
Decr.	1	7 26 00	299 55 31	62 51 30

The longitudes and latitudes of the preceding table, have been deduced from angular distances observed of Arcturus, Vega, Altair,  $\alpha$ ,  $\beta$  and  $\nu$  in the Swan, with the circle of reflection, described in page 265 of this volume.

The observations from the 1st Octr. till the 7th Novr. were made in the city of Havanna, the others at the plantation of Don Joseph de Cotilla, situated in latitude  $22^{\circ}$  55' 16" N. and  $44^{\prime\prime}$ , 3, in time, E. of Havanna.

The times of the observations were determined by a good chronometer, regulated by absolute and corresponding altitudes of the sun and stars, and the times observed at the plantation, are referred to the city of Havanna, by the difference of meridians.

To determine the place of the comet, many series of observations were made with two or three of the above named stars, choosing those that made the most convenient triangles, and as the different observations could not be made at the same time, care has been taken, to refer all the distances observed, to the same instant, by means of the variation observed of the distances of the said stars from the comet.

The distances observed were freed from the effects of refraction, corrected by reference to the state of the thermometer and barometer.

The places of the stars were taken from the Connoisance dc temps, of Paris, 1806; allowance being made for the proper motion, precession of the equinox, nutation and aberration. Further, the latitudes and longitudes of the said table are the apparent, that is, affected by the nutation and aberration. The elements of the orbit of the comet were calculated from the first observations which I made in Havanna, that is, from 1st Octr. to 7th Novr., by Don Francis Leamur, Lieutenant Col. of the Royal Corps of Engineers, and are the following:—

Passage through the perihelion, mean time at	the	city of	Hav	anna,	Septr.	18th 111	• 58'	59''
Longitude of the ascending node.		٠.		8s	26° 3	39′09″		
Inclination of the orbit.				•	63 1	2 30		
Place of the perihelion.				. 9	00 4	15 OI		
Perihelion distance, that of the sun	bei	ng 1.	•	•	0,64	62 <b>128</b>		

After having concluded the observations, namely up to the 1st December, I determined to calculate the elements of the parabolic orbit, by the combination of all the observations, and the following elements are the results.

Passage through the perihelion, mean time, at the city of Havanna, Septr. 18.	12h 37	00‴
at Greenwich,	18 06	40
Longitude of the ascending node from the mean equinox=8' 26° 42' 12"		
Inclination of the orbit 63 12 51		
Place of the perihelion		
Perihelion distance, that of the sun being 1 0,6462667.		

Comparison of the observations with the results of the theory calculated by the above elements.

The longitudes and latitudes observed and calculated in the following table, are freed from nutation and aberration.

The two last columns shew the difference between the longitude and latitude, observed and calculated.

1807.	Mean Hav			The ol Long					served ude.	Calc Long						Diff. long.	
	h	,	"	0	,	W	0	1		0	1	W	0	,			N
Octr. 1	6	54	50	220	21	14	18	46	32 N	220	21	37	18	46	30 N	-23	+02
18	6 3			234			37	41	36	234	36	19	37	42	15	+47	-39
Novr. 3	6 :			251				13		251				12		+ 3	+22
4	64	49	30	252	57	25	51	55	00	252	58	12	51	54	21	-49	+39
7		44		257				54		257				55		+17	-28
17	7 (			272				17		272				18		+13	-48
18	6 9	27	36	274			59	42	49	274				42		-27	09
19	6 4	44	10	276	28	21	60	06	25	276	28	39		06		18	-21
25	6 :	59	07	287	54	41	61	56	37	287	55	14	61	56	51	-33	-14
Decr. 1	7 9	26	00	299	56	18	62	51	30	299	56	44	62	51	22	-26	+08

Continuation of Astronomical Observations, made at the plantation of Don Joseph de Cotilla.

#### Determination of Latitude.

1807, Novr. 1	3 By 8 s circle	eries of ( of reflecti	⊙'s dout ion.	ole al	titud •	es, c 22°	bsei 55	ved 14½	near th ″N.	e me	ridia	n, with a
1	7 di	tto	⊙'s dian	ieter.		22	55	15				
2	1 di	tto	ditto.	•						400	ومربير	10// # >T
	7 By 4 s								mean.	22*	55	13″,5 N
	the	pole-star.	•	•	•	22	55	20				
2	the 0 By 2 s	eries of F	omalhat.			22	55	17				
									mean.	22	55	18, 5
	Me	an Latitu	de	•		•			• •	22	55	16

By astronomical observations, I have determined the bearing of the highest hill of Camoa, N. 13° 34' 10" W.

The hill of Camoa, from the city of Havanna, according to the survey which was made by the order of Government, was determined =29250. Varas of Castilla =13,11 geographical miles, bearing S.  $45^{\circ}$  E.

Latitude of Havanna, according to a great number of obse made with the same circular reflector. Hill of Camoa S. 45 E. 13',11 miles, difference of latitude		•	23°		30″ 16
Latitude of the hill of Camoa. By direct observations on the hill, with the circular.		•	22 22		
Mean latitude of the hill.	•	•	22	59	16

The combination of the two bearings, and the latitudes of the hill of Camoa and Havanna, gives the former E. of the city of Havanna 11' 05'', 2=44'', 3 in time.

Observations made on a lunar eclipse, on the 14th Novr. 1807.

		D / W
The beginning of the eclipse, apparent time.		13 52 12
beginning of immersion of Tycho.	•	14 15 52
end of immersion of Tycho.		14 19 12
beginning of Mare humorum.		14 23 32
end of the eclipse.	•	15 58 42

Observation of apparent lunar distances, observed with the circle of reflection, at the plantation.—The distances in the following table are the result of 4 series of direct and inverse observations.

1807.	•	Appt. time.		Appt. Dist.	Ther.	Barom.
		h / #		0 1 11		
Novr.	14	8 01 12	a 8 C's remote limb	19 06 40	65 <u>1</u>	30 10
		8 26 51	ditto. ditto	18 57 52	-	
		15 37 40	a & C's nearest limb.	16 45 37		
	17	9 07 30	ditto. ditto	20 51 04	66	30 00
		9 24 40	ditto. ditto	21 00 46	7	
	19	20 31 20	⊙ C nearest limbs	118 10 53	72	30 00
		21 17 28	ditto	117 50 13		
	21	21 33 52	ditto.	92 16 37	75	29 96
	22	17 25 27	a mg C's nearest limb	41 43 54	67	30 10
2	24	17 53 15	ditto. ditto	12 37 06	65	30 15
		22 01 38	⊙ <b>⊄</b>	51 56 57	72	30 10
-	~	22 16 12	ditto.	51 51 26		00.00
Decr.	2	22 57 34	ditto	52 58 33	77	30 00
		23 23 09	ditto.	53 09 29 <del>1</del>		
		23 57 20	ditto.	53 22 46		
	3 4	0 50 54	ditto.	53 42 144	74	30 12
	4	3 50 27 4 11 51	ditto.	$\begin{array}{r} 66 & 14 & 57\frac{1}{2} \\ 66 & 19 & 55 \end{array}$	74	30 12
	7	1 51 48	ditto.	99 05 4 <b>1</b>	74	30 10
	6	2 11 58	ditto.	99 03 41 99 12 32	(**	30 10
		6 14 33	C's and Atair nearest limb.	59 12 32	70	29 98
		6 26 30	ditto.	59 00 25	10	23 90
	9	6 33 59	$Ca$ $\gamma$ remote limb.	47 53 01	69	30 00
	15	7 04 51	<b>C</b> a 8 remote limb.	29 03 304	74	29 95
	<b>.</b>	7 16 24	ditto. ditto	29 10 25		~~ ~~
	20	12 13 28	C and Regulus remote limb.	21 25 48	70	30 10
		12 18 16	ditto.	21 28 25		
1808		12 31 02	ditto	21 35 14		
	11	14 06 42	Ca S nearest limb	26 13 05	68	30 15
-		14 23 14	ditto.	26 18 48		
	19	16 08 31	C and Antares nearest limb.	37 44 55	55	30 14
		16 23 02	ditto.	37 39 58		
	21	21 25 35	⊙ <b>⊄</b>	61 55 05	71	30 08
		21 45 24	ditto,	61 49 10		

#### January 11th, 1808. Occultation of , n by the moon.

The disappearance was instantaneous—magnifying power of the telescope, 75.

By 4 series of similar observations on Sirius. 50 36 54,4

By 10 series of angular distances, observed with the circular reflector, and corrected for refraction, the mutual distance was determined  $=36^{\circ}$  17' 19,4".

The difference of right ascension in time, of the above stars =16' 59,5''.

By the distance observed, and the difference of right ascension results the difference of declination.  $36^{\circ}$  08' 00,4''

Taking the latitude of the place as stated above 22° 55' 16" and correcting the meridional altitudes observed, from nutation, aberration, and precession, we have the true, or mean declinations of the two stars on 1st January 1808.

Canopus.		•	•	50°	951	34,9"	
	•	•	•			•	
Sirius.		•	•	16	27	36,8	

Comparing the observations of la Caille on 1750, and supposing the annual precession in longitude=50,1" we have the proper motion of Canopus in declination in 58 years-0' 10,1" Sirius. +1 02,0 Mean declination of Sirius, according to the Rev. Nevill Maskelvne

on the 1st of January, 1808.			•			*	•	- <b>,</b>	•	169	' 27 <b>'</b>	30"
Connoisance de temps. By the observation with the cir	•			•	•		•	•				38,6 36,8
By the observation with the cir	Cuta	rrene	CLOP.		•	•	•		•	10	21	30,0

Astronomical observations made at the city of Havanna. Latitude of the place 23° 08' 30".

Occultations of stars by the moon, observed with an Achromatic telescope—magnifying power 75.

	C 1 a 25 on the dark limb, appar			,		<b>1</b> 1h	<b>5</b> 3′	34″	
	374 of Mayer on the dark limb.		•	•			01		
3d,	w Lion on the dark limb.	ditto.	•	٠	•	10	33	49	
1 ne immers	ions were instantaneous.								

Observations made on a lunar eclipse at the city of Havanna, on the 9th of May, 1808:—magnifying power of the telescope 70.

IMMERSIONS.	Mean time.	EMERSIONS.	Mean time.
Beginning of the Eclipse.	12 22 29	End of total darkness of the C	14 55 10
Beginning of Grimaldus.		End of Grimaldus.	15 01 50
End of ditto	12 27 58		
Beginning of Aristarcus.	12 28 38		
End of ditto	12 30 08	End of Aristarcus	. 15 05 45
Beginning of Mare humorum.	12 37 57	Beginning of Tycho.	15 14 14
ditto. of Copernicus	12 39 37	End of ditto	15 15 47
End of ditto	12 40 57	Center of Schikardus.	. 15 19 28
Beginning of Plato	12 43 57		
End of do	12 45 07	End of Plato	15 21 23
Beginning of Mare serenitat.	12 50 57	Beginning of Mare serenitat.	
Center of ditto	12 55 16	Center of ditto	. 15 33 07
Beginning of Tycho		End of ditto	. 15 37 57
End of ditto	12 58 41	End of Taruntius.	. 15 44 17
Beginning of Mare Crisium.	13 09 21		. 15 46 46
End of ditto.	13 14 20		15 49 26
End of Langrenus	13 18 00		•
Total darkness of the C.	13 21 25	The end of the eclipse.	. 15 54 36

The above observations of the lunar eclipse are very exact, excepting the beginning and the end of the eclipse, which are liable to the error of one and a half minute, on account of the strong penumbra.

Table of the results of the occultations of the stars by the moon.

	at pl	anta	<b>n C</b> ation 1808.	Ha	war	ma.	Ha	van	na.	Ha	van	na.
	h	1		h	,	"	h	,	"	h	1	4
Mcan time of immersions.	14	54	32	11	56	07	8	58	14	10	31	26
Longitude west from Paris, .	5	38	06	5	38	50	5	38	50	5	38	50
Mean time at Paris.			38			57			04		10	
Apparent longitude of the stars.			55			03,6			10,5			
Apparent latitude of the stars.	-		48 S.	-		34 S.	5	20	48 S,	5	34	07 S.
Latitude—Vertical angle.			52			13						
Logarithmic radius of the earth.			036			000		~~	4.0		20	05 9
Equatorial horiz. parallax of the C Parallax in Longitude.			23,8 36,0			24,5			15			05,3 52,3
Parallax in latitude.			05,2			35,9 17,9			<b>44,5</b> 38,3			16,2
Apparent difference of latitude be-		-10	03,2		-23	17,9	+	- 44	30,3	+	.01	10,2
tween the moon and stars.		5	03		6	08,0		7	36,0		14	31
Conjunction mean time.	13		34		-	••,•			00,0		••	
Havanna west from the plantation.			44.3				1					
Conjunction in Havanna by observ.			49,7	11	h09	37	8	18	11	9	34	32
At Paris by the new tables.	19	38	24,0	16	48	51	13	57	11	15	13	40
TT		00		-		-						
Havanna west from Paris.	1 2	39	34,3	5	S9	14	5	39	00	5	39	08

#### Results of observed lunar distances.

	January 11th, 1808 January 19th, 1808. Cay Cay K Antares. C & Antares.
	h ' W h / W h / W h ' W
Apparent time of the observations. Apparent distances nearest limb.	14   06   43   14   23   18   16   08   31   16   23   02     26°13   05   26°18   48,2   37   44   45   37   39   55
Altitudes of $C$ calculated $\begin{cases} Appt. \\ True. \end{cases}$	43 03 40 39 27 20 47 18 20 48 57 00   43 49 31 40 10 30 47 57 34 49 35 01,5
Altitudes of the stars do. {Appt. True.	16   51   40   13   07   10   13   33   30   16   13   20     16   48   41   13   03   20   13   29   41   16   10   09
Corrected distances Apparent longitude of the stars. Apparent latitude of ditto.	27 12 22,8 27 21 06,4 38 37 06,1 38 29 02,8   67 06 52,3 247 05 00,6 5 28 47 32 30 5   5 28 47 S. 4 32 30 S. 5 30 5
True longitude of the moon by observations January 11th, 14 <sup>h</sup> 15' 00' Apparent time at the Plantation.	er-} 3° 04° 21′ 02″,7
January 19th at 16 <sup>h</sup> 15' 46",5 appa	rent time 6 28 30 03,5
Longitude of the Plantation W. from Ditto from the observation of 1	
Solar eclipse of June	16th, 1806, in the city of Havanna.
Apparent time.	Dist. of the horns.
8 55 34,6 beginning of the   8 57 20,2 . .   8 59 22,0 . . .   9 02 08,6 . . .   9 04 35,8 . . .   9 07 44,0 . . .   9 11 40,0 . . .	e eclipse 0 00,0 6 12,9 8 51,6 Observed by Don Antonio de 11 40,0 > Robredo, with a Heliometer 13 31,5 15 17,0 17 19,3
the conjunction, by the begi By the first observation of dist By the second.	ance
By the third	225104
Ditto. in Paris, page 296	
Havanna west from Paris.	· · · · <u>5 39 09</u>

## By the Solar Eclipse (page 162,) observed in the city of Havanna,

# and at Lancaster in Pennsylvania. U. S.

Havanna west from Lancaster Lancaster west from Paris (p	197.)		••••	•.	•••	•	-	24 14	25″ 41	
Havanna west from Paris.	•	•	•		•	•	5	39	06	

#### ASTRONOMICAL OBSERVATIONS

# Longitude of Havanna, by the observations compared with the new tables published at Paris in 1806.

#### January 11, 1808. 5 39 34 Occultations of stars. April 5. 5 39 14 May 2. 5 39 00 May 3. 5 39 08 $\begin{cases} \alpha & \forall \text{ January 11,} \\ \alpha & \forall \text{ January 19.} \end{cases}$ 5 39 14 Distances of moon. 5 39 03 Solar eclipse, 1803. 5 38 16 1806. 5 38 20 do. Moon's eclipse, May 9, 1808. 5 38 51 5 38 55 By corresponding observations of solar eclipse February 21, 1803 5 39 06 June 16, 1806. 5 39 09 Ditto 5 39 07 Havanna inferred from Philadelphia, by the chronometer, No. 63. 5 39 18 5 38 37 Inferred from Veracruz, page 225. Ditto from Porto Rico, page 225. 5 38 34 5 38 50 Havanna west from Paris. 5 38 57

#### Passage of Venus over the disk of the Sun, June 3d, 1769.

Elements from Astronomical tables at			10h 11'	47" mean time at Paris.
Longitude of the sun, apparent equinox.		•	73° 27	18,3
Right ascension of the sun.			72 03	16
Horary motion in $\bigcirc$ 's right ascension.			2	34
Relative horary motion in longitude .			3	57,40
Horary motion of Venus in latitude S.			0	35,42
Inclination of the orbit			8 29	10,00
Apparent obliquity of the ecliptic .			23 28	11,5
Radius vector of the earth.	•	•	1,01519	90
Radius vector of Venus	•		0,72626	50
O's semidiameter.			15' 47",	07

By a previous calculation of the observations of this passage, I had determined the following elements:—

Sun's parallax at the mean distance from the earth =	8~,62378
Apparent conjunction, mean time at Paris = 10 <sup>h</sup> 11' 4	
Apparent longitude of Venus	
Duration of the passage between the interior contacts $= 5^{h} 4$	1′ 54″,5 in mean time.
or 5 4	1 52,1 in apparent time.
Latitude of Venus at conjunction, north 1(	) 15,94
	09,18
	5 15,89
	5 13,27
Difference of Venus and sun's parallaxes at the passage $=$	21,352

#### TABLE I.

#### Reduction of the observations to the center of the earth.

		of	he	ob-	0	f	Appt. time of con. cen- ter of earth		from						mer	r of earth. id of Paris IV.					
		h	,	"	,	H	h	,	"	h	,	H				h	,	"	h	,	#
Petersburg.	III	15	24	41	5	16	15	19	25	-1	51	56				13	27	29			
	IV			27	-4	58	15		29										13	46	33
Cajaneburg.	II			45	+6	44	9			1	41	47	7	45	48						
11	IV			27	-4	36	15		51		ير بر	~	L		~~				13	46	10
Wardhus.	II			10 24	$+6 \\ -4$	27	15		37 5 <b>1</b>	-1	55	07	7	45	30		27	44			
				24 41	$  -4^{4} -4$				32							13	21	44	12	46	95
Batavia.	III			13	-4				11	6	58	15				13	27	56	10	-10	20
Data Ha.	iv			31	_3				46			10				1	21		13	46	31
Gurief.	Î			25	6		16	45	57	-3	18	24	1			13	27	33			
	IV	17	11	06	6	06	17	05	00										13	46	36
Oremburg.	III	17	05	06					24	-3	30	58				13	27	<b>5</b> 6			
U	IV			24					31										13	46	33
Orsk.	ш			26					17	3	44	43	L			13	27	34			
	IV			57					05		00	~~					~~	~	13	46	22
Pekin.	III			24						-7	30	30				13	27	27	10	10	20
	[ IV	] 21	20	54	1-3	54	121	23	00	ł			I			1			μ3	46	30
			I	Alear	1 rest	ilts	of th	e I	II a	nd I	v c	Cont	ac	ts	•	13	27	39,9	916	46	27,5

In the calculation of this and the following tables, the parallax of the sun, at the mean distance of the earth =8''62378, and the difference of parallaxes at the passage =21'',352.

Note. The III contact at Petersburg was observed 13<sup>h</sup> 28' 29" and I subtracted one minute of time, being probably an error committed in setting down the time of the clock.

#### TABLE II.

Reduction of the observation to the center of the earth.

		Apparent time of observations.	Effect of parallax.	Appt. time at the center of the carth.	Longitudes	
Paris. Greenwich. Kew. Oxford. London.	11	<b>b</b> , <b>w</b> 7 38 45 7 29 25 7 28 17 7 24 20 7 29 16	, $+7 \ 03,1$ +7 04,2 +7 04,2 +7 02,0 +7 02,0 +7 04,0	h / //////////////////////////////////	$\begin{array}{c cccc} h & \prime & \bullet \\ +00 & 09 & 21 \\ + & 10 & 24 \\ + & 14 & 23 \\ + & 9 & 37 \end{array}$	h / # 7 45 48,1 7 45 50,2 7 45 45,2 7 45 45 7 45 57
Stockholm. Upsal.	ean.	8 41 46 8 40 12	$\begin{array}{r} +6 56,0 \\ +6 57,4 \\ \hline +7 01,6 \end{array}$	8 48 42 8 47 09	- 1 02 55 - I 01 15 Mean.	7 45 47 7 45 54,4 7 45 49,5

	Apparent time observations.	Effect of parallax.	Appt. time at the center of the earth h / W
Fort Prince of Wales III IV	7 19 20	+4 12,1 +0 39,1 +0 49,5	1 19 35,1 7 01 26,1 7 20 09,5
St. Joseph III IV	6 13 19	+0 20,3 +4 47,9 +4 46,0	0 17 47,3 5 59 37,9 6 18 05,0 21 38 30,6
Taity $\begin{cases} II \\ III \end{cases}$	21 44 04 3 14 08	-5 33,4 +6 17,4	3 20 25,4
Philadelphia $\begin{cases} I \\ II \end{cases}$	2 13 45 2 31 28	+3 38 +3 54	2 17 23 2 35 22
Cape Francais Cambridge II	2 26 12 2 44 44,5 2 47 30,0	+2 23,6 +2 37,6 +4 19,0	2 28 35,6 2 47 22,1 2 51 49,0

#### TABLE III.

#### TABLE IV.

Difference of time between the interior and exterior contacts at the

#### center of the earth.

Petersburg 19 04   Wardhus. 18 41   Batavia. 18 35   Oremburg 18 37   Gurief. 19 03   Fort Prince of Wales. 18 42   St. Joseph. 18 27   Greenwich. 18 48   Cape Francais. 18 47   Ingress. 18 41
Mean 18 46,4
b/IV contact at Paris, center of the earth, Table I.13 46 27,5 effect of parallax.Mean result of Table IV18 46,4
III contact by the observations of IV contact.13 27 41,1-4 54,1By the mean of direct observations, Table I.13 27 39,9-5 18,1
Mean result for the III contact.   <
Total duration of the interior contacts (n) $5 41 51,0$ $-12 07,7$ By the observations of Wardhus. $5 42 14",0$ $5 41 54,8$ $-11 10,0$ By dittoCajaneburg. $5 41 35 5$ $5 41 54,8$ $-11 10,0$
Mean.(a) $5$ $41$ $52,9$ $-11$ $38,8$ By the observations at Taity $5$ $41$ $54,8$ $+11$ $50,8$ By the observations at St. Joseph $5$ $41$ $50,6$ $+$ $4$ $27,6$ Byditto.F. P. Wales $5$ $41$ $51,0$ $ 3$ $33,0$

#### Results of sun's parallax at the mean distance of the earth.

By the duration, at Taity and (n) Taity and (a)					8′′,600 8,620
Taity and Wardhus	87,7312		• •		8,623
St. Joseph and F. P. Wales.	• •		•		8,623
Taity and F. P. Wales. = $\cdot$	8,645 <b>}</b> 8,588 <b>}</b>	•	·	•	8,616
Mean result	• •	• •	•	•	8,615

Contacts at the center of the earth, for the meridian of Paris; allowing the sun's parallax at the mean distance of the earth=8",615.

																h	'		
IC	ontac	t.	Ap	par	ent	tim	e.		•	•						7	27	02,5	
11				•			•	•		•	•		•		٠	7	45	48,9	
111	•	•			•				•			•				13	27	41,4	
IV				•											•	13	46	27,8	
Error of the	durat	ion	of	the	obs	erv	atio	ns a								•		•	+22",8
									C	ajan	leb	urg			•			•	-16,0
									- (1	n)		-	•		•		•	•	- 0,5
									Ť	`aity		•		•					+ 1,3
										t. Íc					•	•		•	- 2,3
									F	· P.	W	ale	s.		•		•	•	- 1,7

Determination of the longitude of different places, from Paris, by the observation of the passage of Venus.

	:	ь <i>/</i>	Ħ	h	,	#	
Philadelphia, by the $\begin{cases} I \\ II \end{cases}$	CALCITOR COMMENCE	5 10	40,0 27,5	<u> </u>	10	03,7	w.
Cape Francais $\begin{cases} I \\ II \end{cases}$			27,5	4	58	27,5	w.
Cambridge, N. Eng. II				4	54	00,5	w.
Taity $\begin{cases} II \\ III \end{cases}$	$\cdot \cdot $		17,9 16,6 }	10	07	17,2	w.
St. Joseph $\begin{cases} II \\ III \end{cases}$			01.6	7	28	02,8	W.
F. P. Wales. $\cdot  \begin{cases} II\\III \end{cases}$		526 523	13,5 15,3	6	26	14,4	w.
Wardhus. $\begin{cases} II \\ III \\ IV \end{cases}$	• • • •	L 54 L 55 L 55	47,6 09,9 07,1	1	55	01,5	E.
Cajaneburg $\begin{cases} II \\ IV \end{cases}$	• • • •	l 41 l 41	39,67 23,65	1	41	31,5	E.
Gurief $\begin{cases} III \\ IV \end{cases}$		318 318	32 5	3	18	23,5	E.
Oremburg $\begin{cases} III \\ IV \end{cases}$		3 30 3 31		3	30	52,5	E.
Orsk $\begin{cases} III \\ IV \end{cases}$		344 344	35 }	3	44	36,0	E.
Batavia	• • • •	658 658		6	58	24,0	E.
Pekin $\begin{cases} III \\ IV \end{cases}$		736 736	$16$ $\overline{33}$	7	86	24,5	E.
Petersburg {III IV		1 51	44 01 }	1	51	52,5	E.

#### Passage of Mercury over the disk of the Sun, Novr. 12th, 1782.

		h 🕫		
Philadelphia.		9 34 9 40 10 51 10 57	$\begin{array}{c} 00\\ 30\\ 35 \end{array}$ Mean time.	
Paris	$\begin{cases} 1 & \cdot & \cdot \\ 11 & \cdot & \cdot \\ 111 & \cdot & \cdot \end{cases}$	2 58 2 04 • 4 17	04,5 30 40 Apparent time.	
Greenwich	и	2 54	42 Apparent time.	
Cambridge in New England.	$\begin{cases} III & \cdot & \cdot \\ IIII & \cdot & \cdot \\ IV & \cdot & \cdot \end{cases}$	. 10 12 11 23 11 29	10 06 14 Apparent time.	
Difference of $\odot$ and $\heartsuit'$	s semidiameters.	• • •	. 16' 04'',27	
Difference of horizontal Horary relative motion i		• •	· 4, 01 · 3 53, 45	
Horary motion of $\breve{a}$ in l	atitude, N. 🛛 .		. 51 91	
Appt. conjunction at Paris, Apparent conjunction, by of Apparent conjunction Longitude of Philadelph Cambridge	observations at P	hiladelphi Cambridge ris.	nia 225	4′ 0 <u>9</u> ′′ 3 59 .0 16

Passage of Mercury over the disk of the sun, Novr. 5th, 1787.

#### Observations.

Apparent time
Paris, interior contact at the ingress
Viviers
Cadiz do 0 44 30
Marseilles do 1 31 07
Montauban
Vienna
Prague do
I 20 08 00
Philadelphia. $\begin{cases} II & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 20 & 09 & 30 \\ III & \cdot & \cdot & \cdot & \cdot & \cdot & \cdot & 0 & 0 & 0 & 0$
IV 1 01 14
$C^{I}$
Cambridge in XII
N. England. III 1 15 44
IV
Montevideo III
Difference of the horizontal parallaxes. $4^{\prime\prime}$ ,149
Horary relative motion in longitude between the ingress and conj. 349,55
Between the egress and conjunction
a diameter of ⊙—1",50 irradiation
Apparent conjunction at Paris, by the observations in Europe = 3 33 16
Philadelphia.
Cambridge
Montevideo
Longitude of Philadelphia west from Paris
Cambridge west from Paris 4 53 40
Montevideo west from Paris 3 54 15

#### Annular eclipse, April 3d, 1791.

# Elements from the Astronomical tables published at Paris, in the year 1806, by order of the Commissioners of longitude.

1791. April 3. Astronomical mean time at Paris.0 54 40©'s longitude from the apparent equinox.13°41 58©'s right ascension in time.0h 50 25©'s semidiameter.0°16 00,42Equation of time.+ 3 17,53©'s horary motion in longitude.2 27,59Horary motion in O's right ascension in time.9,10Horary diminution of the equation of time.0,90C's longitude from the apparent equinox.13°41 7,83C's longitude from the apparent equinox.9,10Horary diminution of the equation of time.0,90C's longitude from the apparent equinox.13°41 7,8C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.86Apparent obliquity of the celiptic.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude.00,75Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in latitude.40,11		1	1	N
$\bigcirc$ 's longitude from the apparent equinox. $13^{\circ}4158$ $\bigcirc$ 's right ascension in time. $0^{h}5025$ $\bigcirc$ 's semidiameter. $0^{\circ}1600,42$ Equation of time. $+317,53$ $\bigcirc$ 's horary motion in longitude. $227,59$ Horary motion in $\bigcirc$ 's right ascension in time $9,10$ Horary diminution of the equation of time. $0,90$ $\circlearrowright$ 's longitude from the apparent equinox. $13^{\circ}4137,8$ $\circlearrowright$ 's equatorial horizontal parallax. $5436,1$ $\bigcirc$ 's equatorial horizontal parallax. $86$ Apparent obliquity of the ecliptic. $232753,0$ Moon's horary motion in longitude. $246,72$ Horary diminution of $\circlearrowright$ 's horizontal parallax. $246,72$ Horary diminution of $\circlearrowright$ 's horizontal parallax. $0,075$ Equation of 2d order of the $\circlearrowright$ 's horary motion in longitude. $-0,040$ dittodittoin latitude. $+00,111$	1791. April 3. Astronomical mean time at Paris.	0 5	4	40
O's right ascension in time.0h 50 25O's semidiameter.0°16 00,42Equation of time.+ 3 17,53O's horary motion in longitude.2 27,59Horary diminution of the equation of time.9,10Horary diminution of the equation of time.0,90C's longitude from the apparent equinox.13°41 37,8C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.8,6Apparent obliquity of the ecliptic.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.0,75Equation of 2d order of the C's horary motion in longitude				
<b>O'16</b> 00,42Equation of time. <b>C</b> 's horary motion in longitude. <b>Horary motion in O's right ascension in timeHorary motion in O's right ascension in timeHorary diminution of the equation of time</b> . <b>Horary diminution of the equation of time</b> . <b>C's longitude from the apparent equinos</b> . <b>C's north polar distance</b> . <b>C's equatorial horizontal parallax</b> . <b>O'16</b> 00,42 <b>C's equatorial horizontal parallax</b> . <b>C's equatorial horizontal parallax</b> . <b>C's equatorial horizontal parallax</b> . <b>C's horary motion in longitude</b> . <b>Moon's horary motion in longitude</b> . <b>Moon's horary motion in latitude</b> S. <b>Horary diminution of C's horizontal parallax</b> . <b>C's horizontal parallax</b> .				
Equation of time.+ 3 17,53O's horary motion in longitude.2 27,59Horary motion in O's right ascension in time9,10Horary diminution of the equation of time.0,90C's longitude from the apparent equinos.13°41 37,8C's north polar distance.89 15 05,9C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.30 12,97Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax				
$\bigcirc$ 's horary motion in longitude.2 27,59Horary motion in $\bigcirc$ 's right ascension in time9,10Horary diminution of the equation of time.0,90 $\circlearrowright$ 's longitude from the apparent equinox.13°41 37,8 $\heartsuit$ 's north polar distance.89 15 05,9 $\heartsuit$ 's equatorial horizontal parallax.54 36,1 $\circlearrowright$ 's equatorial horizontal parallax.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of $\heartsuit$ 's horizontal parallax.00,75Equation of 2d order of the $\heartsuit$ 's horary motion in longitude				
Horary motion in O's right ascension in time910Horary diminution of the equation of time.0,90C's longitude from the apparent equinox.13°41 37,8C's north polar distance.89 15 05,9C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.8,6Apparent obliquity of the ecliptic.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude				
Horary diminution of the equation of time.0,90C's longitude from the apparent equinox.13°41 37,8C's north polar distance.89 15 05,9C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.8,6Apparent obliquity of the ecliptic.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.0,75Equation of 2d order of the C's horary motion in longitude			~	
C's longitude from the apparent equinox.13°41 37,8C's north polar distance.89 15 05,9C's equatorial horizontal parallax.54 36,1O's equatorial horizontal parallax.8,6Apparent obliquity of the ecliptic.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude				
C's north polar distance. 89 15 05,9   C's equatorial horizontal parallax. 54 36,1   O's equatorial horizontal parallax. 8,6   Apparent obliquity of the ecliptic. 23 27 53,0   Moon's horary motion in longitude. 30 12,97   Moon's horary motion in latitude S. 246,72   Horary diminution of C's horizontal parallax. 00,75   Equation of 2d order of the C's horary motion in longitude. -00,40   ditto ditto in latitude. + 00,11	C's longitude from the apparent equinor.	304	11	
C's equatorial horizontal parallax. 54 36,1   O's equatorial horizontal parallax. 8,6   Apparent obliquity of the ecliptic. 23 27 53,0   Moon's horary motion in longitude. 30 12,97   Moon's horary motion in latitude S. 2 46,72   Horary diminution of C's horizontal parallax. 00,75   Equation of 2d order of the C's horary motion in longitude. -00,40   ditto ditto in latitude. + 00,11	Cs north polar distance			
O's equatorial horizontal parallax.8,6Apparent obliquity of the ecliptic.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude 00,40dittodittoin latitude.				
Apparent obliquity of the ecliptic.23 27 53,0Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude		Ŭ		
Moon's horary motion in longitude.30 12,97Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude.00,40dittodittoin latitude.+ 00,11		2 9	7	
Moon's horary motion in latitude S.2 46,72Horary diminution of C's horizontal parallax.00,75Equation of 2d order of the C's horary motion in longitude 00,40dittodittoin latitude.+ 00,11	Moon's horary motion in longitude			
Horary diminution of C's horizontal parallax. Equation of 2d order of the C's horary motion in longitude. ditto ditto ditto in latitude. + 00,11				
Equation of 2d order of the C's horary motion in longitude — 00,40 ditto ditto in latitude — 00,11				
ditto ditto ditto in latitude + 00,11	Foustion of 2d order of the C's horary motion in longitude			
	ditto ditto ditto in latitude.	-		
Proportion of the acustonial boniz name and the Ce haviz diameter 60, 99 151	Proportion of the equatorial horiz, paral, and the C's horiz, diameter. 60			
Proportion of the equatorial and polar diameters of the earth = 330 : 329			24	40,1

Observations made by the Rev. Nevil Maskelyne, at Greenwich.

0h 18' 40" Apparent time, beginning	of the ccl	ipse.	
1 44 51 Least distance of the limb	os. 12′52	<i>ii</i>	
3 06 47 End of the eclipse.			
By the mean result of 8 observations, O's diame	ter was	. 31′ .	57″,0
		-	
h	/ #	<u></u> ከ / //	h / #
Apparent time of the observations at Greenwich. 0 1	8 40	1 44 51	3 06 47
	4 27,8	54 26,7	54 25,8
Parallax in longitude	8 02,4	-29 07,0	
	4 47,1	-30 18,6	-27 10,4
C's apparent semidiameter-2' inflexion. 1	5 02,2	15 01,0	14 59,0
()'s semidiameter-2" irradiation 1	5 58,4	15 58,4	15 58,4
Conjunction at Greenwich by the combination of the be	eginning a	nd	-
the end of the eclipse apparent time	me.		0h 45′ 16″,5
Correction of latitude by the tables.	•	• •	+ 13
By the least distance of the limbs.	• •	• •	+ 13,6
Supposing the irradiation of the sun's semidiameter	•	= 1″,8	•
The O's diameter was observed	1′57″,0		
By the tables	2 00,8		
The corrected distance of the limbs = $\frac{32' 00'',8 \times 12}{31}$	52" 19	5011 5	
The corrected distance of the minos $=$ 31,57	14	JJ ,J	
The double irradiation	• —	3, 6	
True distance of the limbs.		49, 9	- 1 1111 1

And the correction of moon's latitude corrected from the effect of refraction = $+11^{\prime\prime},5$ Conjunction at Paris = $(0^{h} 45' 16^{\prime\prime},5+9' 21^{\prime\prime})=00^{h} 54' 37^{\prime\prime},5$ 

Observations at the National Observatory of Paris.

Beginning of the ecl	ipse,	apparent	t time.	•	•	•	= 0h	361	55″	,4
End of the eclipse.	•••	•	•	۰	•		3	20	52	Q

Conj. at Paris by the	combination o	of the beg	ginning s	ind e	nd of	the ecl	ipse.	0h	54'	38''.5
Correction of the C's	latitude by th	ne new ta	ables	•	•	· .	·	+	00	09
Palermo, Steginning, end.	, apparent tin	ne.	• •	•	•	• •	-	2	10	17, 0
		• •	•	•	•	• .		3	59	20,4
Conjunction at Palerr	no		• •	•	•	•		1	18	45
Conjunction at Palerr Conjunction at Paris		1h 38′ 4:	5",5	-44	′ 06″	3	2	0	54	39, 5
Correction of C's latit	ude .	•	•	•	•	•		+		11, 0
Petersburg { apparent	t time, begin	ning of th	he eclips	e.	•	• •	•	2	56	30
receisburg 2	end of	f the ecli	pse.		•		•	5	21	
By the combination of	the beginnin	ig and en	d, conju	nctio	on in r	nean tir	ne ==	2	46	37, 6
Conjunction at Paris,	mean time	= (2h	46' 37'	,6	1h 5	i' 56")	-	0	54	41, 6
Correction of the C's	latitude by tl	he tables			•	• • •	-	+		10, 0
	-									-
			1							v
By the observations o	f Greenwich,						ion of	f C's	lat.	=+11,5
By ditto	Paris.		. (	54	38,5	•	•	•	•	+ 9,0
By ditto	Palermo.	•	. 0	54	39,5		•		•	+11
By ditto	Petersburg	• •	. (	54	41,0	•	•	•	•	<b>÷</b> 10
-	-									-
Conjunction at Paris,				54	39	•		•	•	+10,3
Correction of C's long	ritude by the	new tab	les.	•	•	• •	•	•	,	=+20,3

## Observations at Cambridge, New England.

											-			
	h	,												
April 2	18	01	27		AD	pare	ent	t time, beg	rinni	ing	of the ed	linse		
		08						formation.		0				
		12						break.						
		28						ne eclipse.						
					h	. ,	,		ħ	,		b	,	
Apparent time of obser	vatio	n.			1	9 08	8	07	19	12	56	20	28	26
Moon's latitude by tab			<b>.</b> 3	N		47	7 9	21.7		47	08,4			38,7
C's equatorial horizon								28.1			28,1			27.2
Parallax in longitude.					•	2	1	46,4			34,6			36,1
Parallax in latitude.	•							28,6			18,3			52,6
Apparent latitude of the	ie C	1.1	s.			0	0	06,9			09,9			13,9
Horizontal & diameter		he (	Σ	•	•			54,32			54,32			54,23
Augmentation of the @				er.	•			4,00			4,27			7,22
<b>C</b> 's apparent semidiam	eters.		•			1	4	58,32		14	58,59		15	01,45
O's semidiameter from	n the	tab	les.			1	6	00,42		16	00,42			00,42
Difference and sum of	semi	dia	met	ers			1	02,10		1	01,83		31	01,87
Horary relative motion annular and the ti							e	formation	of tł	ne	• •	27	<b>4</b>	5″,8

alimitar and the time of the conjunction					•	- 21	45,8
Between the end of the eclipse and the conj	unction.	• •				. 27	45.2
Results · difference of semidiameter betwee	n the form	ation on	d the				,-
breaking of the annular, by observation.	• •	•		•	•		61,45
breaking of the annular, by observation. By the Tables. $\frac{1'\ 02'',10+1'\ 01'',83}{2}$	• •	•	•	•	•	•	61,96
Correction of the difference of semidiameter Correction of the sum of semidiameters.							
					h	,	
Conjunction from the annular for						00 4	0,8
annular bro	eaking.	•	•		20	00 4	0,8
end of the	eclipse.	• •	•		20	00 4	0,8
Longitude west from Paris .		•	<u>4</u> h	53'	58".9	,	

# Observation in the City of Philadelphia.

Formation of annulus. 18 46 11.5 Apparent time	
Break of annulus 18 50 28,5 Observed by Mr. Rittenhous	se
Formation of annulus. 18 46 11,5 Apparent time   Break of annulus. 18 50 28,5   End of the eclipse. 20 03 42   h h	
h = h = h = h = h = h	
Apparent time of the observation. 18 46 11,5 18 50 28,5 20 03 42   C's latitude by the tables + 10",3 N. 00 47 37,5 00 47 25,5 44 01   Parallax in latitude. - - 47 07,1 46 59,3 43 58   Apparent latitude of the C N 00 30,4 00 26,2 00 03   Parallax in longitude. - 24 35,3 24 28,3 20 33   C's apparent semidiameter. 14 57,35 14 57,56 15 00   Semidiameter of the sun. - 16 00,42 16 00,42 16 00   Diff. and sum of C and O's semidiameters. 1 03,07 1 02,86 31 01   With the corrections $-0''.5$ for the difference of semidiameters and $-4''.4$ for the semidiameter 50 00 50 00	s I O
Parallax in latitude. $-4707.14659.34358$	3.5
Apparent latitude of the C . N . 00 30,4 00 26,2 00 03	3,4
Parallax in longitude	5,4
U's apparent semidiameter	),74
Diff. and sum of $\mathbb{C}$ and $\bigcirc$ 's semidiameters. 1 03.07 1 02.86 31 01	16
of semidiameters, according to the results of the observations at Cambridge, we have	the
following results :	
By the breaking of the annulus $1944437$ By the breaking of the annulus $194438$ 19h 44' 37",	6
By the end of the eclipse. $194438$	
Conj. by the formation of the annulus. Mean time. 19h 44' 37" By the breaking of the annulus 19 44 38 By the end of the eclipse 19 44 38 Longitude of Philadelphia west from Paris = 5 10 01,	4
Observations at George Town, Maryland.	
C C	
h / W	
Formation of annulus.18 36 43Apparent timeBreak of annulus.18 39 57.By the end of the eclipse.19 52 21.	-
By the end of the eclipse, 19 52 21	lsq.
•••••••••••••••••••••••••••••••••••••••	
Conjunct. by the formation of annulus, mean time.193700By the breaking of ditto.193700By the end of the eclipse.1936'58",5Longitude of George Town west from Paris.51740, 5	
By the breaking of ditto. $19 37 00 \ge 19h 36' 58'', 5$	
Longitude of George Town west from Paris = 5 17 40 5	
Note. I have subtracted I' of time from the formation and the breaking of the annul	lus.
from the observations at Philadelphia, and added 1' of time to the formation of the annu	
at George Town, those errors having been discovered by the result of the observations.	lus
De the combination of the charmentions of the complemention of the	lus
By the combination of the observations of the annular eclipse of the sun, April 3, 17	ılus 791,
By the combination of the observations of the annular eclipse of the sun, April 3, 17	ılus 791,
By the combination of the observations of the annular eclipse of the sun, April 3, 17	ılus 791,
By the combination of the observations of the annular eclipse of the sun, April 3, 17	ılus 791,
By the combination of the observations of the annular eclipse of the sun, April 3, 17	ılus 791,
By the combination of the observations of the annular eclipse of the sun, April 3, 17 I have determined the corrections of the Irradiation of the $\bigcirc$ 's semidiameter $=$ 1",70 inflex. of $\bigcirc$ 's semidiameter $=$ 2" Page 298 of this Volume. 1806. Total eclipse of the $\bigcirc$ - 1, 87 1764. Annular eclipse of the $\bigcirc$ - 2, 15 1801. Occultation of $\alpha$ mp $\bigcirc$ 1799. Passage of $\clubsuit$ over the $\bigcirc$ -1, 50	ilus 91, 93 35 82
By the combination of the observations of the annular eclipse of the sun, April 3, 17 I have determined the corrections of the Irradiation of the $\bigcirc$ 's semidiameter $=$ 1",70 inflex. of $\bigcirc$ 's semidiameter $=$ 2" Page 298 $\begin{cases} 1806. \text{ Total eclipse of the } \bigcirc -1, 87 & \cdots & -1, \\ 1764. \text{ Annular eclipse of the } \bigcirc -2, 15 & \cdots & -1, \\ 1801. \text{ Occultation of } \alpha \text{ m} \bigcirc \\ 1799. \text{ Passage of } \emptyset \text{ over the } \bigcirc -1, 50 & \cdots & -1, \end{cases}$ Mean correction of the irradiation $\ldots -1, 80$ inflexion. $\ldots -1, $	ilus 91, 93 35 82 75
By the combination of the observations of the annular eclipse of the sun, April 3, 17 I have determined the corrections of the Irradiation of the $\bigcirc$ 's semidiameter $=$ 1",70 inflex. of C's semidiameter $=$ 2" Page 298 $\begin{cases} 1806. \text{ Total eclipse of the } \bigcirc -1, 87 & \cdots & -1, \\ 1764. \text{ Annular eclipse of the } \bigcirc -2, 15 & \cdots & -1, \\ 1801. \text{ Occultation of } z = 10 & \bigcirc -1, 50 & \cdots & -1, \\ 1799. \text{ Passage of } \heartsuit \text{ over the } \bigcirc -1, 50 & \cdots & -1, \\ \text{Mean correction of the irradiation} & -1, 80 & \text{ inflexion.} & -1, \\ Recapitulation of the results of longitudes of Philadelphia and Call$	ilus 91, 93 35 82 75
By the combination of the observations of the annular eclipse of the sun, April 3, 17 I have determined the corrections of the Irradiation of the $\bigcirc$ 's semidiameter $=$ 1",70 inflex. of $\bigcirc$ 's semidiameter $=$ 2" Page 298 $\begin{cases} 1806. \text{ Total eclipse of the } \bigcirc -1, 87 & \cdots & -1, \\ 1764. \text{ Annular eclipse of the } \bigcirc -2, 15 & \cdots & -1, \\ 1801. \text{ Occultation of } \alpha \text{ m} \bigcirc \\ 1799. \text{ Passage of } \emptyset \text{ over the } \bigcirc -1, 50 & \cdots & -1, \end{cases}$ Mean correction of the irradiation $\ldots -1, 80$ inflexion. $\ldots -1, $	ilus 91, 93 35 82 75
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